

Energy transport in gravity waves

let us consider a gravity wave propagating along the x -direction in an ideal fluid, within an incompressible potential flow approximation. The system is y -invariant, z vertical axis, and we know

$$\left. \begin{aligned} \varphi(x, z, t) &= A \cosh[k(z-h)] \cos(\pi x - \omega t) \\ \zeta(x, t) &= -\frac{1}{g} \omega A \cosh(kh) \sin(\pi x - \omega t) \end{aligned} \right\} \text{for a wave in a basin with bottom } z=h \text{ and vacuum (negligible fluid) on top}$$

Let us consider a surface through which we want to calculate the energy flux. This is a plane with normal unit vector $\hat{\mathbf{e}}_x$, width L in the y -direction, and height measured between the bottom $z=h$ to the top of the fluid layer $z=\zeta$; let us call this surface $S(\zeta)$. When we studied the energy flux we found out the flux density is $\rho(E_m + \omega)\bar{v}$, that is here $\rho\left(\frac{1}{2}v^2 + gz + p/\rho + \epsilon\right)\bar{v}$ and by ignoring ϵ (conserved in an incompressible flow) the flux through $S(\zeta)$ results

$$Q^E(S) = \int_S \rho\left(\frac{1}{2}v^2 + gz + p/\rho\right) \bar{v} \cdot d\bar{S} = \rho \int_S \left(\frac{1}{2}v^2 + gz + p/\rho\right) v_x dy dz$$

Making use of the generalized Bernoulli's equation applied between a generic z and ζ ,

$$\frac{1}{2}v^2 + \frac{p}{\rho} + gz + \frac{\partial \varphi}{\partial t} = \frac{1}{2}v^2(\zeta) + p_0 + g\zeta + \frac{\partial \varphi(\zeta)}{\partial t},$$

pressure at the free surface is a constant which we can set as $p_0 = \phi$ (negligible for very light upper fluid), the kinetic terms are quadratic, hence negligible in a linear approximation; finally, the interface boundary condition yields

$$g\zeta(x, t) + \frac{\partial}{\partial t}\varphi(x, \zeta, t) = \phi \quad \text{and therefore we simplify the equation above into}$$

$$\begin{aligned} \frac{1}{2}v^2 + \frac{p}{\rho} + gz &= -\frac{\partial \varphi}{\partial t} \\ \Rightarrow Q^E(S) &= - \int_y^{y+L} \int_{z=-h}^{z=\zeta} \rho \frac{\partial \varphi}{\partial t} v_x dy dz = -\rho L \int_{z=-h}^{z=\zeta} \frac{\partial \varphi}{\partial t} v_x dz = -\rho L \int_{-h}^{\zeta} \frac{\partial \varphi}{\partial t} \frac{\partial \varphi}{\partial x} dz \end{aligned}$$

Now let us split the integral over two subintervals, $(-h, \phi]$ and $[\phi, \zeta]$; the second term, discarding higher-order infinitesimal contribution, can be approximated using the value in $z=\phi$ of the integrand, so that

$$Q^E(S) = -\rho L \int_{-h}^{\phi} \frac{\partial \varphi}{\partial t} \frac{\partial \varphi}{\partial x} dz - \rho L \frac{\partial \varphi(x, \phi, t)}{\partial t} \frac{\partial \varphi(x, \phi, t)}{\partial x} \zeta(x, t)$$

The first term is second order in magnitude, while the second one is a third-order contribution. Furthermore, we are interested in calculating a period-average quantity, and one could see that the second term, as a product of three sinusoidal function, has a zero average. So let us make a time average over a period T :

$$\langle Q^E(S) \rangle = -\rho L \frac{1}{T} \int_{t_0}^{t_0+T} \int_{z=-h}^{z=0} \frac{\partial \phi}{\partial t} \frac{\partial \phi}{\partial x} dz dt \quad \text{and plugging in the explicit form of } \phi$$

$$\frac{\partial \phi}{\partial t} \frac{\partial \phi}{\partial x} = \omega A \cosh[k(z+h)] \sin(kx - \omega t) \cdot (-k) A \cosh[k(z+h)] \sin(kx - \omega t) = \\ = -k \omega A^2 \cosh^2[k(z+h)] \sin^2(kx - \omega t)$$

$$\text{and since } \langle \sin(kx - \omega t) \rangle = \frac{1}{2}, \quad \int \cosh^2(\alpha x) dx = \frac{1}{4\alpha} \sinh(2\alpha x) + C$$

$$\langle Q^E(S) \rangle = \frac{1}{2} \rho L k \omega A^2 \int_{-h}^0 \cosh^2[k(z+h)] dz = \frac{1}{2} \rho L k \omega A^2 \left[\frac{1}{4k} \sinh(2k(z+h)) + \frac{z+h}{2} \right]_{-h}^0 = \\ = \frac{1}{8} \rho L \omega A^2 [2kh + \sinh(2kh)]$$

Now let us calculate the time-averaged amount of mechanical energy stored in a region $R(\zeta)$ that is just upstream of $S(\zeta)$ and extends in the x direction of a quantity $\Delta x = \lambda$ (one wavelength). Any point within $R(\zeta)$ is a point (x, y, z) such that

$$\begin{cases} x_0 < x < x_0 + \lambda \\ y_0 < y < y_0 + L \\ -h < z < \zeta(x, t) \end{cases}$$

We also call $R(\phi)$ a subregion of $R(\zeta)$ of parallelepipedal shape obtained considering only the part with $z \in [-h, \phi]$. The boundary surface of $R(\phi)$ shall be called $S_R(\phi)$ and the top surface of $S_R(\phi)$ (points $(x, y, z=\phi)$) shall be $A(\phi)$.

The mechanical energy in $R(\zeta)$ is now written subtracting from it the energy in the absence of the wave (when the region occupied by the fluid at rest is just $R(\phi)$):

$$E_m(R(\zeta)) = \int_{R(\zeta)} \rho \left(\frac{1}{2} V^2 + g z \right) d^3x - \int_{R(\phi)} \rho g z d^3x$$

In the first term, the region above $z=\phi$ contributes an amount of energy that is approximated (neglecting higher-order infinitesimal terms) with the integrand value in $z=\phi$, then

integrating in x and y over $A(\phi)$ and multiplying by $\zeta(x,t)$; the potential energy contribution over the region $R(\phi)$ is also subtracted and we get

$$\bar{E}_m(R(\zeta)) = \int_{R(\phi)} p \frac{t}{2} v^2 d^3x + \int_{A(\phi)} p \frac{t}{2} v^2 \zeta(x,t) dx dy + \underbrace{\int_{A(\phi)} \left(\int_{z=\phi}^{z=\zeta} pg dz \right) dx dy}_{\rightarrow \frac{1}{2} \zeta^2(x,t)}$$

once again the second term is not only negligible as a higher-order infinitesimal, but made of oscillating functions whose overall result is a vanishing time average, so

$$\bar{E}_m(R(\zeta)) = \frac{1}{2} p \int_{R(\phi)} v^2 d^3x + \frac{1}{2} pg \int_{A(\phi)} \zeta^2(x,t) dx dy$$

We shall now use some algebra to manipulate the first integral; indeed $v^2 = |\text{grad}\phi|^2$

$$\text{and we can notice that } \text{div}(\phi \text{grad}\phi) = \text{grad}\phi \cdot \text{grad}\phi + \phi \nabla^2 \phi = |\text{grad}\phi|^2 = v^2$$

$$\begin{aligned} \Rightarrow \bar{E}_m(R(\zeta)) &= \frac{1}{2} p \int_{R(\phi)} \text{div}(\phi \text{grad}\phi) d^3x + \frac{1}{2} pg \int_{A(\phi)} \zeta^2(x,t) dx dy = \\ &= \frac{1}{2} p \int_{S_R(\phi)} \phi \text{grad}\phi \cdot \hat{n} da + \frac{1}{2} pg \int_{A(\phi)} \zeta^2 dx dy \end{aligned}$$

We can cut out some parts of the flux integral over $S_R(\phi)$, and precisely where $\hat{v} = \text{grad}\phi$ has a zero normal component with respect to the integration surface, or where the net flux is zero:

- At the bottom, $z=-h$, the b.c. requires $v_z = \text{grad}\phi \cdot \hat{e}_z = \phi$;
- Due to y -invariance, $v_y = \text{constant}$ and thus on the two surfaces $y=y_0$, $y=y_0+L$ of the volume $R(\phi)$ the incoming and outgoing fluxes are equal and balance out;
- Periodicity along x requires incoming and outgoing fluxes through the surfaces at $x=x_0$, $x=x_0+L$ to be equal and cancel out again.

What is left is just the flux through $A(\phi)$ in the vertical direction; so

$$\bar{E}_m(R(\zeta)) = \frac{1}{2} p \int_{A(\phi)} \phi(x,y,t) \frac{\partial \phi}{\partial z} \Big|_{z=\phi} dx dy + \frac{1}{2} pg \int_{A(\phi)} \zeta^2(x,t) dx dy =$$

$$\text{recalling } \zeta = -\frac{1}{g} \frac{\partial \phi}{\partial t} \Big|_{z=\phi}$$

$$= \frac{1}{2} \rho \int_{A(\phi)} \left[\phi(x, \phi, t) \frac{\partial \phi}{\partial z} \Big|_{z=0} + \frac{1}{g} \left(\frac{\partial \phi}{\partial t} \right)^2 \Big|_{z=0} \right] dx dy$$

Let us rewrite explicitly ϕ and its derivatives before we perform the time average:

$$\phi(x, z, t) = A \cosh(k(2t\hbar)) \cos(kx - \omega t) \rightarrow \phi(x, \phi, t) = A \cosh(k\hbar) \cos(kx - \omega t)$$

$$\frac{\partial \phi}{\partial z} \Big|_{z=0} = Ak \sinh(k\hbar) \cos(kx - \omega t)$$

$$\frac{\partial \phi}{\partial t} \Big|_{z=0} = Aw \cosh(k\hbar) \sin(kx - \omega t)$$

$$\textcircled{1} \quad \frac{1}{T} \int_T \phi(x, \phi, t) \frac{\partial \phi}{\partial z} \Big|_{z=0} dt = \frac{1}{T} \int_T A^2 k \sinh(k\hbar) \cosh(k\hbar) \cos^2(kx - \omega t) dt = \frac{1}{4} A^2 k \sinh(2k\hbar)$$

$$\Leftrightarrow = \frac{1}{2} \sinh(2k\hbar)$$

$$\textcircled{2} \quad \frac{1}{T} \int_T \frac{1}{2} A^2 w^2 \cosh^2(k\hbar) \sin^2(kx - \omega t) dt = \frac{1}{2} A^2 w^2 \cosh^2(k\hbar) = \text{using } \omega^2 = g/k \tanh(k\hbar)$$

$$= \frac{1}{2} A^2 k \sinh(k\hbar) \cosh(k\hbar) = \frac{1}{4} A^2 k \sinh(2k\hbar)$$

We find here that the two terms are equal. This should not really come as a surprise: It is a more general fact that for small oscillations, the time averages of kinetic and potential energy are the same. So we finally end up writing

$$\langle \bar{E}_m(R(\zeta)) \rangle = \frac{1}{2} \rho \int_{A(\phi)} \frac{1}{2} A^2 k \sinh(2k\hbar) dx dy = \frac{1}{4} \rho A^2 k \sinh(2k\hbar) \cdot \lambda L \quad \text{with } \lambda = 2\pi/\hbar$$

$$\Rightarrow \langle \bar{E}_m(R(\zeta)) \rangle = \frac{1}{2} \bar{\rho} \lambda L A^2 \sinh(2k\hbar)$$

In general terms, we can write the flux of energy multiplying the energy density per unit volume ρ_{en} times the flux velocity v_f and the area \sum of the cross section S through which we calculate the flux; inverting the expression, we can get v_f as follows:

$$\langle Q^E(S) \rangle = \rho_{en} \sum v_f = \rho_{en} \underbrace{\frac{(\sum)}{\lambda}}_{\text{volume}} v_f = \frac{\langle \bar{E}_m(R(\zeta)) \rangle}{\lambda} v_f$$

$$\Rightarrow v_f = \lambda \langle Q^E(S) \rangle / \langle \bar{E}_m(R(\zeta)) \rangle = \lambda \frac{1}{8} \rho L \omega A^2 [2k\hbar + \sinh(2k\hbar)] / \frac{1}{2} \bar{\rho} \lambda L A^2 \sinh(2k\hbar) =$$

$$= \frac{\lambda \omega}{4\bar{\rho}} \left[1 + \frac{2k\hbar}{\sinh(2k\hbar)} \right] = \frac{\omega}{ek} \left[1 + \frac{2k\hbar}{\sinh(2k\hbar)} \right] = v_g$$

$v_f \approx v_g$ group velocity tells us that energy is not propagated at phase velocity, but at group velocity (even for a monochromatic wave*). Thus for gravity waves the deepest physical meaning of the group velocity is that of mean energy transport velocity (even more than that of translation velocity of a wave packet, which is always undergoing a progressive deformation in a dispersive medium).

[* = The carrier wavelength will be that of the wave itself.]

Appendix - Phase and group velocity (a very brief summary)

A monochromatic (single-wavelength) wave featuring an angular frequency ω and wavenumber k propagates with a velocity called PHASE VELOCITY $v_p = \omega/k$ (with $k=2\pi/\lambda$, λ wavelength). This ω/k ratio does not necessarily have to be constant; on the contrary, v_p and ω can be a function of k . The expression $\omega(k)/k$ is called DISPERSION RELATION (the name comes from the fact that propagation in a "dispersive medium" will yield different velocities for waves with different wavelength/wavenumber, hence a "spreading" of a wavepacket made out of waves with a range of k values).

We can say that the dispersion relation is

NORMAL: v_p is an increasing function of increasing λ

ANOMALOUS: v_p is a decreasing function of increasing λ

TRIVIAL: $v_p = \omega/k = \text{constant}$ (ω is a constant; see, e.g., d'Alembert's wave equation, for electromagnetic waves in vacuum)

Whatever dispersion relation we have, a monochromatic wave with angular frequency ω and wavenumber k can be expressed as

$$f(x, t) = A \exp[i(kx - \omega t)] A \exp[ik(x - v_p t)] = g(x - v_p t)$$

for a simple, one-dimensional case; a more general solution to a wave equation is a whole superposition of monochromatic waves; in a continuous limit, this is an integral

$$F(x, t) = \int_{-\infty}^{+\infty} f(k) \exp[i(kx - \omega(k)t)] dk \quad \text{i.e. a } \underline{\text{wavepacket}}$$

with $f(k)$ amplitude (a function that must be such that the integral is acceptable). A physically meaningful and interesting case is a superposition of waves within a limited range of k , i.e. $f(k) = 0$ outside a certain interval of k values and a peak amplitude within that range corresponding to a value $k = k_0$. The wave corresponding to such k_0 has maximum amplitude and is called CARRIER WAVE of the wavepacket. We now want to evaluate the propagation velocity of the wavepacket as a whole.

Since $f(k) \neq 0$ only over a limited k range, $\omega(k)$ can be expanded to first order:

$$\omega(k) \approx \omega(k_0) + \frac{d\omega}{dk} \Big|_{k=k_0} (k - k_0) = \omega(k_0) + \omega'_0 (k - k_0)$$

where we called $\omega'_0 = \frac{d\omega}{dk} \Big|_{k=k_0}$

$$\begin{aligned}
\Rightarrow F(x,t) &= \int_{-\infty}^{+\infty} f(k) \exp[i(kx - \omega(k)t)] dk = \int_{k_1}^{k_2} f(k) \exp[i(kx - \omega(k)t)] dk = \\
&\quad (\text{where } [k_1, k_2] \text{ interval} / f(k) = \phi \quad \forall k \notin [k_1, k_2]; \quad k_1 \in [k_1, k_2], \quad f(k_0) = \max(f(k))) \\
&= \int_{k_1}^{k_2} f(k) \exp\left\{i\left[kx - (w_0 + (k-k_0)\omega'_0)t\right]\right\} dk = \\
&\quad kx - k_0 x + k_0 x - w_0 t - (k-k_0)\omega'_0 t = (k-k_0)(x - \omega'_0 t) + k_0 x - w_0 t \\
&= \int_{k_1}^{k_2} f(k) \exp\left\{i\left[(k-k_0)(x - \omega'_0 t) + k_0 x - w_0 t\right]\right\} dk = \rightarrow \text{a constant with respect to } k \\
&= \exp[i(k_0 x - w_0 t)] \int_{k_1}^{k_2} f(k) \exp\left\{i\left[(k-k_0)(x - \omega'_0 t)\right]\right\} dk = \quad \text{with } \eta = k - k_0 \\
&= \exp[i(k_0 x - w_0 t)] \int_{k_1 + \eta_0}^{k_2 + \eta_0} f(\eta) \exp[i\eta(x - \omega'_0 t)] d\eta = \\
&= \exp[i(k_0 x - w_0 t)] g(x - \omega'_0 t)
\end{aligned}$$

which represents a carrier wave $\exp[i(k_0 x - w_0 t)]$ carrying the wavepacket with an amplitude $g(x - \omega'_0 t)$; the translation velocity of this amplitude is ω'_0 , and we call it

$$\boxed{v_g = \omega'_0 = \frac{d\omega}{dk} \Big|_{k=k_0}} \quad \text{GROUP VELOCITY}$$

The propagation is not "painless": If in the expansion we had considered higher-order terms, in the calculation we would have found a slow deformation (dispersion) of the wavepacket. In other words, the wavepacket does not simply undergo a rigid translation.

The occurrence of a normal/anomalous dispersion relation impacts the value of v_g with respect to that of v_{ph} :

$$\frac{dv_{ph}}{dk} = \frac{d(\omega)}{dk} = \frac{1}{k^2} \left(\frac{d\omega}{dk} - \omega \right) = \frac{1}{k^2} (v_g k - \omega) \gtrless \phi$$

$$\Leftrightarrow v_g k - \omega \gtrless \phi, \quad \text{i.e. } v_g \gtrless \omega/k = v_{ph}$$

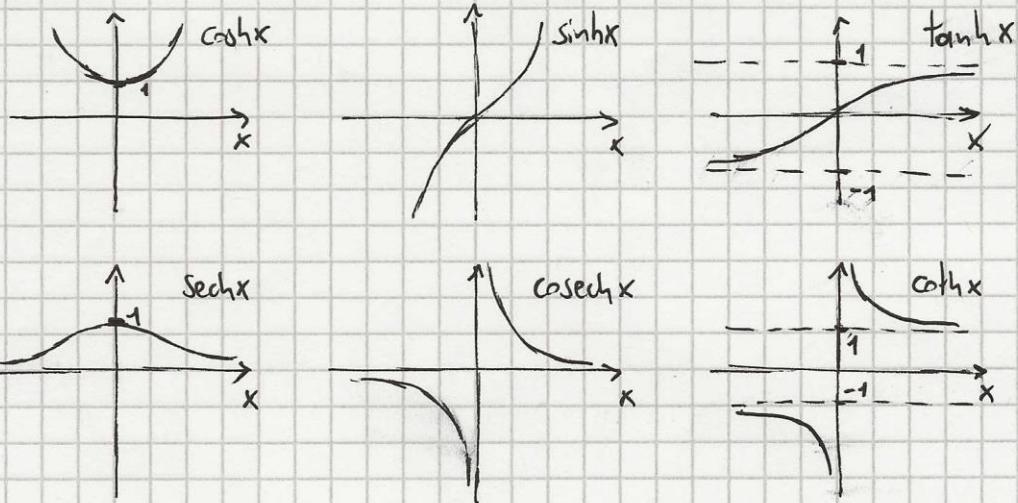
$$\text{so } v_g < v_{ph} \Leftrightarrow \frac{dv_{ph}}{dk} < \phi \quad \text{i.e. } \frac{dv_{ph}}{d\lambda} > \phi \quad \text{normal dispersion relation}$$

$$v_g > v_{ph} \Leftrightarrow \frac{dv_{ph}}{dk} > \phi \quad \text{i.e. } \frac{dv_{ph}}{d\lambda} < \phi \quad \text{anomalous dispersion relation}$$

Appendix - A digest of hyperbolic functions

$$\cosh x = \frac{e^x + e^{-x}}{2}; \sinh x = \frac{e^x - e^{-x}}{2}; \tanh x = \frac{\sinh x}{\cosh x} = \frac{e^x - e^{-x}}{e^x + e^{-x}}$$

$$\operatorname{sech} x = 1/\cosh x; \operatorname{cosech} x = 1/\sinh x \quad \forall x \neq 0; \coth x = 1/\tanh x \quad \forall x \neq 0$$

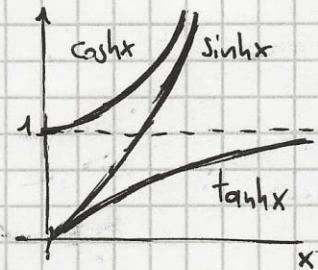


Properties

$$\cosh^2 x - \sinh^2 x = 1$$

$$\frac{d}{dx}(\cosh x) = \sinh x; \frac{d}{dx}(\sinh x) = \cosh x$$

$$\frac{d}{dx}(\tanh x) = \frac{d}{dx}\left(\frac{\sinh x}{\cosh x}\right) = \frac{\cosh^2 x - \sinh^2 x}{\cosh^2 x} = \frac{1}{\cosh^2 x} = \operatorname{sech}^2 x$$



$$\text{Double angle formulae: } \cosh(2x) = \sinh^2 x + \cosh^2 x = 2\sinh^2 x + 1 = 2\cosh^2 x - 1.$$

$$\sinh(2x) = 2\sinh x \cosh x$$

$$\tanh(2x) = 2\tanh x / (1 + \tanh^2 x)$$

Taylor expansion for $x \rightarrow 0$

$$\sinh x \approx x + \frac{x^3}{3!} + \frac{x^5}{5!} + \dots = \sum_{n=0}^{+\infty} \frac{x^{2n+1}}{(2n+1)!}$$

$$\cosh x \approx 1 + \frac{x^2}{2!} + \frac{x^4}{4!} + \dots = \sum_{n=0}^{+\infty} \frac{x^{2n}}{(2n)!}$$

$$\tanh x \approx x - \frac{x^3}{3} + \frac{2x^5}{15} + \dots$$

$$\coth x \approx \frac{1}{x} + \frac{x}{3} - \frac{x^3}{25} + \dots$$

Integrals

$$\int \sinh^2(dx) dx = \frac{1}{4} \sinh(2dx) - \frac{x}{2} + C$$

$$\int \cosh^2(dx) dx = \frac{1}{4} \sinh(2dx) + \frac{x}{2} + C$$